



Tau Neutrino Oscillation & IceCube

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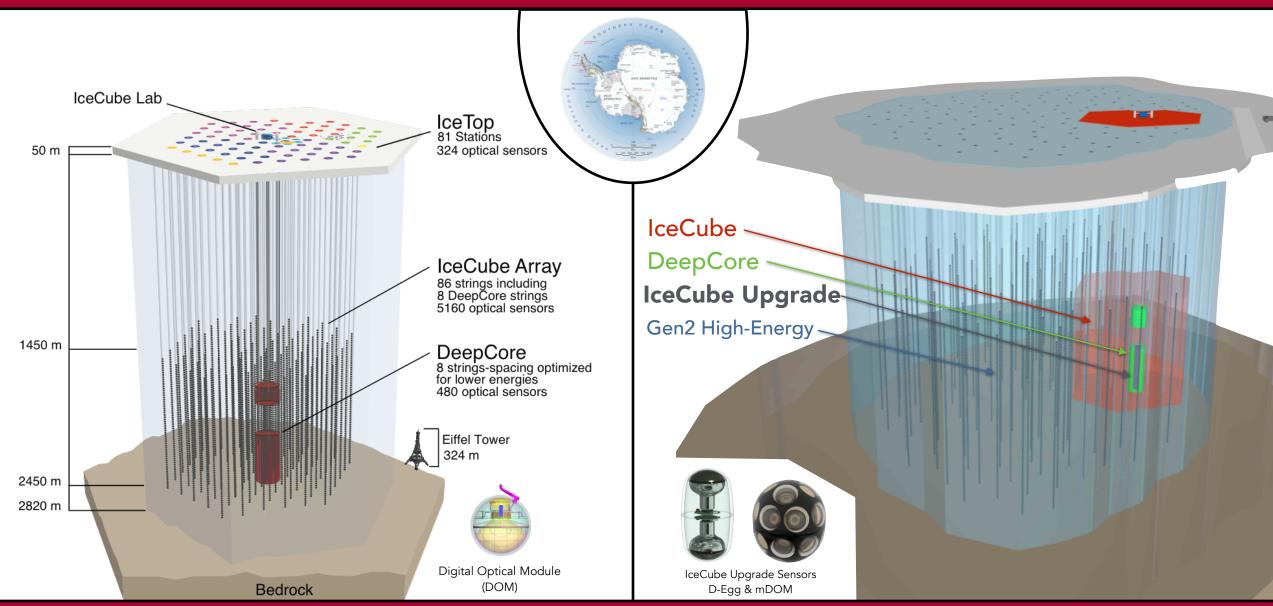
Workshop on Tau Neutrinos from GeV to EeV (NuTau2021) September 2021





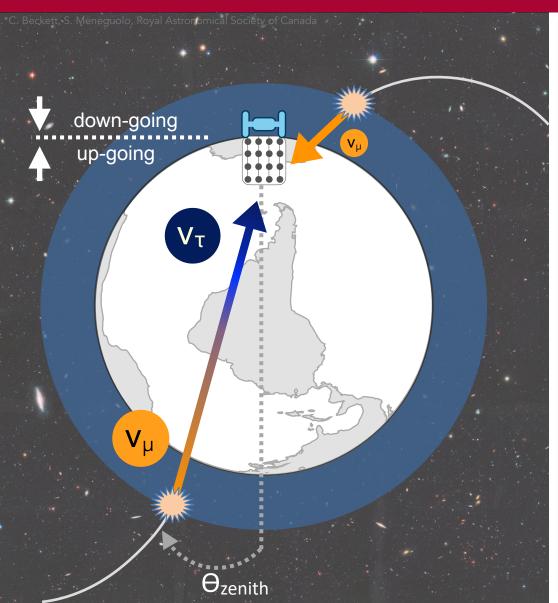
Detectors

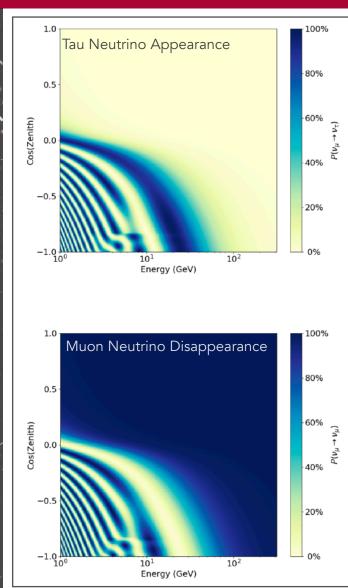


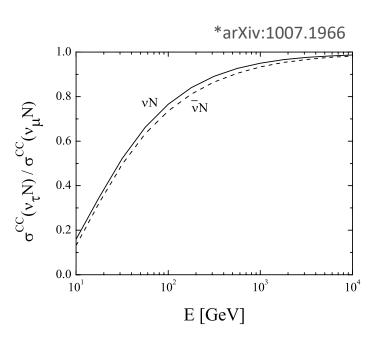


COPENHAGEN Why IceCube/DeepCore/Upgrade for Atmospheric ν_{τ} ?







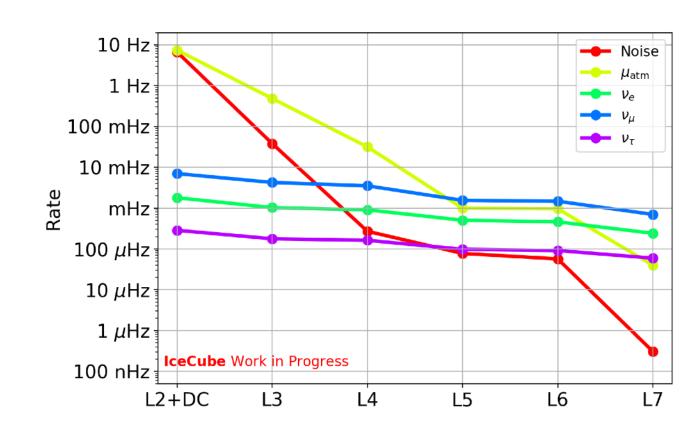




Background & Rates



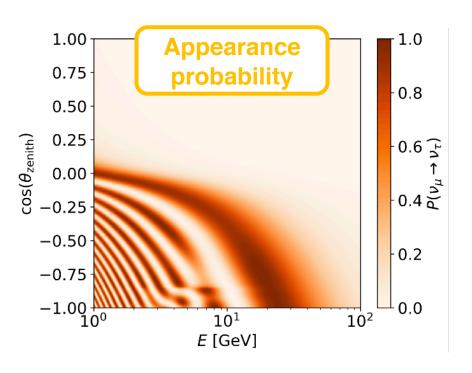
- Iterative approach to event selection
 - Backgrounds reduced by 5-7 orders of magnitude*
 - Detector/electronics pure-noise
 - Atmospheric muons
 - Combination of straight cuts and boosted decision trees
 - Due to high event rates, the high fidelity reconstructions are introduced near end of the selection chain



*upcoming 8-year high-statistics event selection

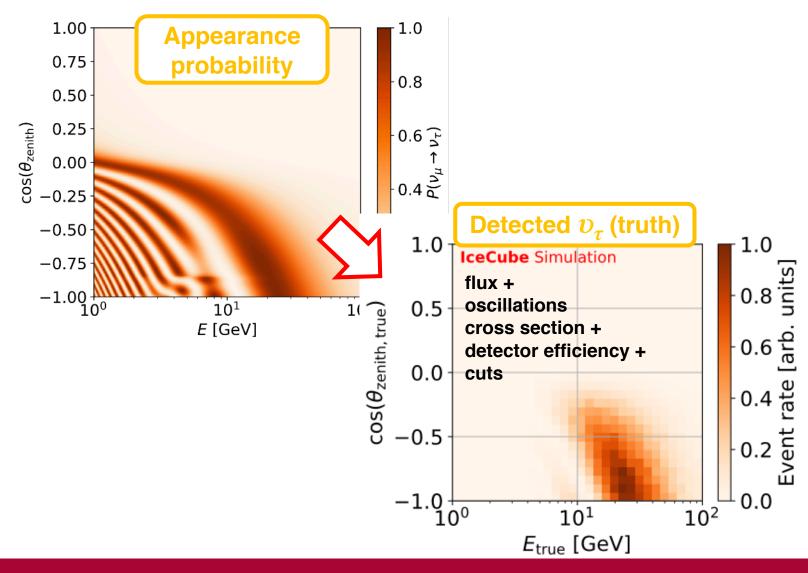
ν_{τ} appearance @ DeepCore





$u_{ au}$ appearance @ DeepCore

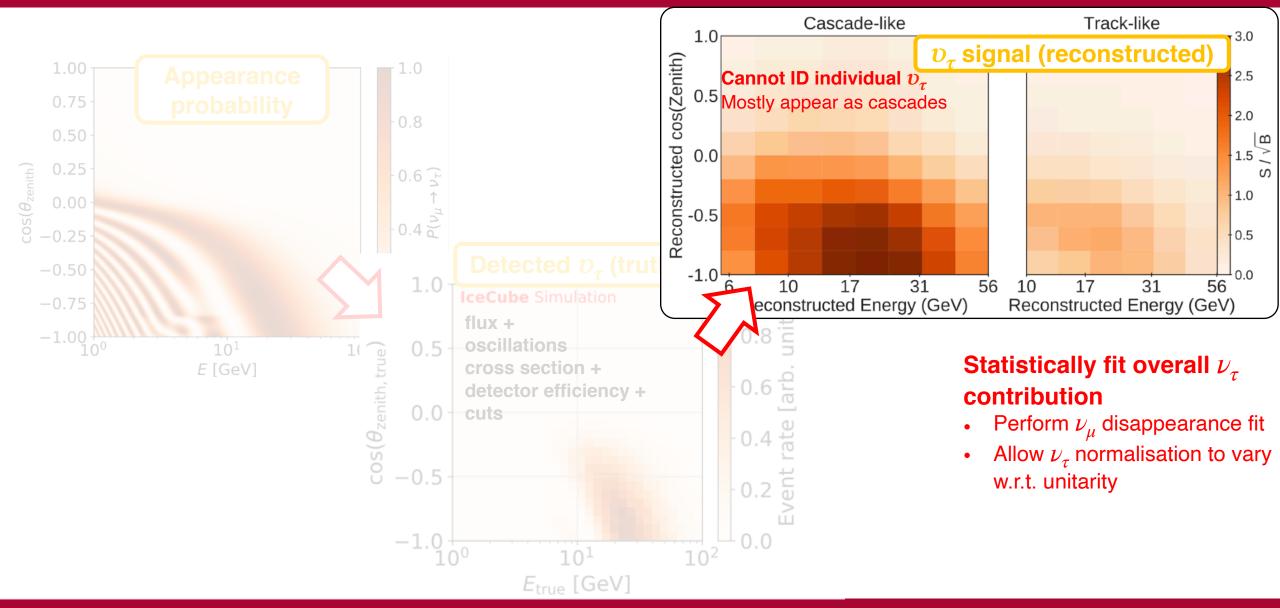






ν_{τ} appearance @ DeepCore



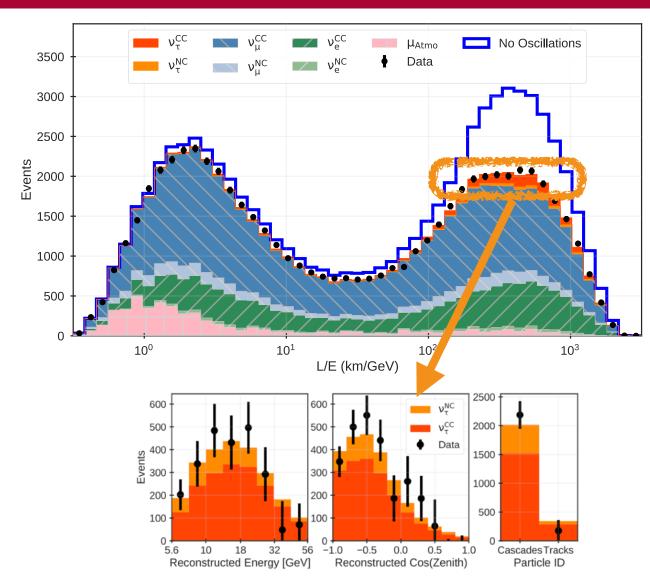




L/E 'Signal'

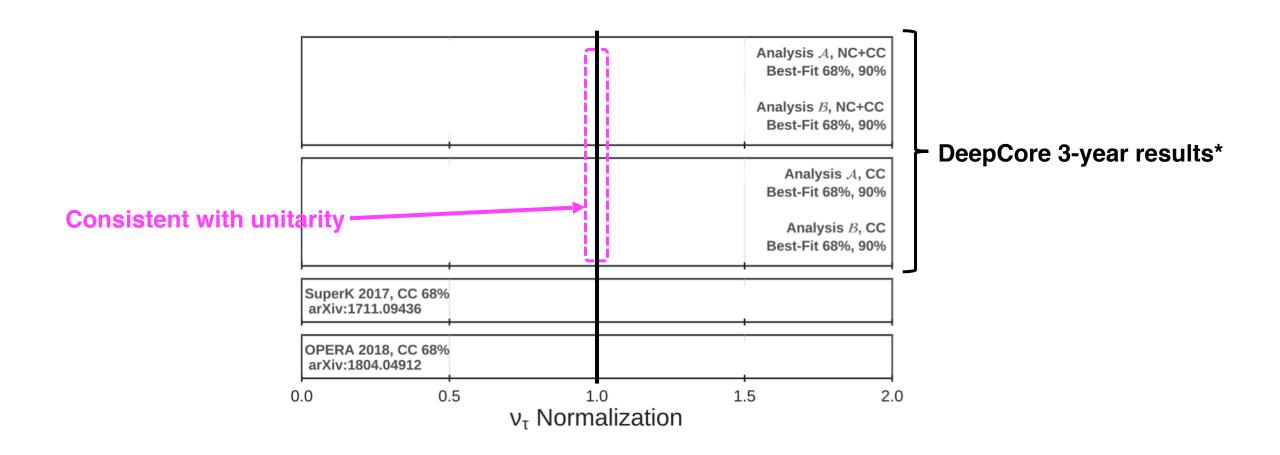


- ν_{τ} appearance is inclusive/ statistical analysis
 - 3-year dataset (2018*)
 - ~60k events, ~1.8k nutau CC
 - April 2012 May 2015
 - 5.6 GeV 56 GeV
 - 8-year dataset (soon)
 - ~250k events, ~10.8k nutau CC [prediction]
 - 5 GeV 300 GeV
- Analyzed in 3 dimensions
 - Particle identification bins
 - 3-year used cascades & tracks
 - 8-year uses cascades, mixed, & tracks
 - Energy
 - Zenith (for baseline)



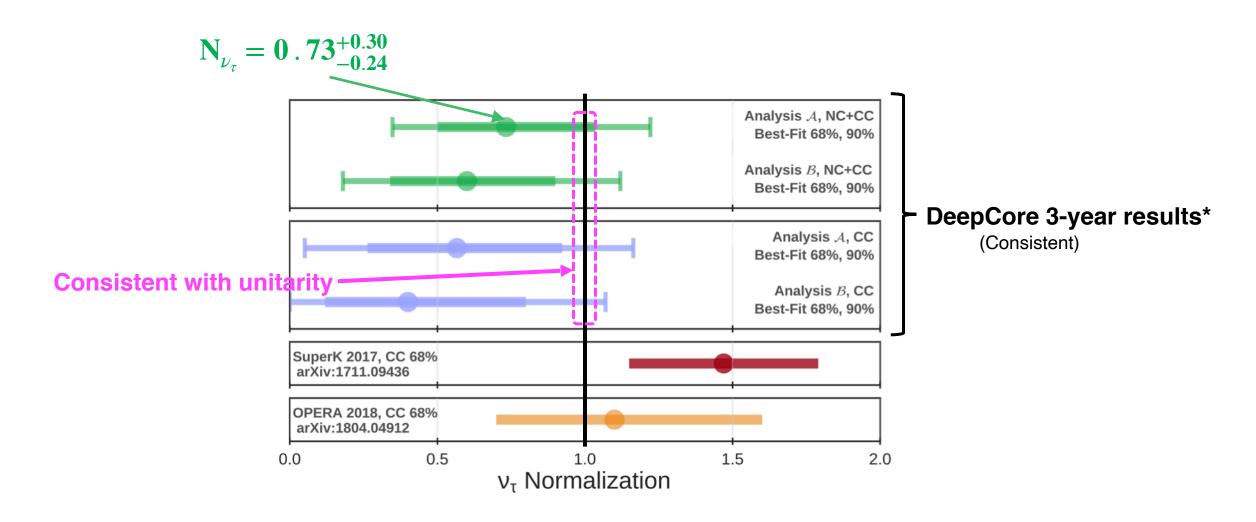
Results





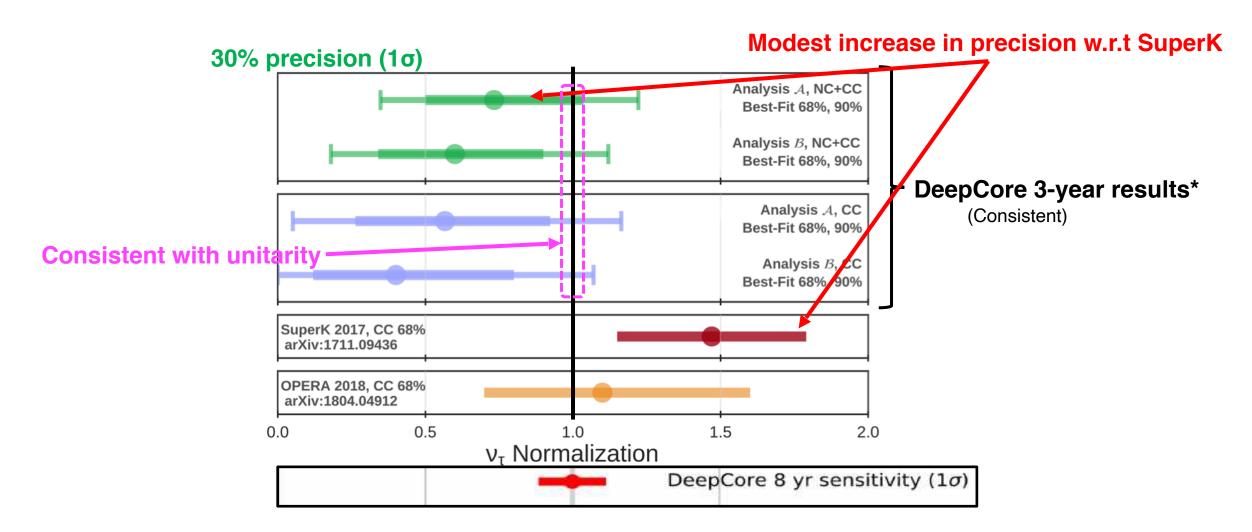
Results





Results







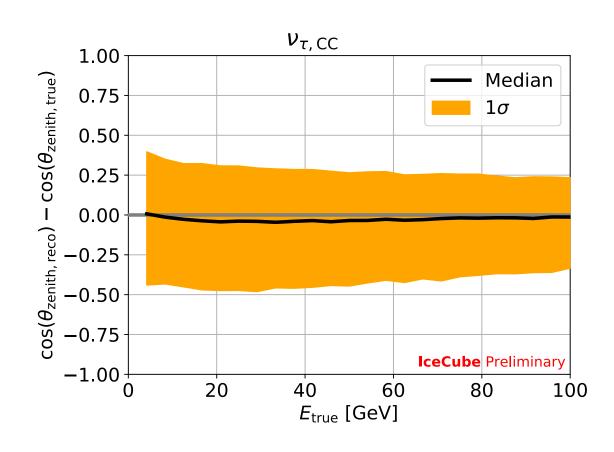


EXPERIMENTAL TECHNIQUES & FEATURES

Cascade Reconstructions (ν_{τ})



- Previous analyses forward light propagation
- Upcoming analysis backwards light propagation
- Future analysis & IceCube Upgrade likelihood free inference, convolution neural network, graph neural network
 - Faster reconstructions will provide better event selections... positive feedback loop
 - Better reconstruction resolutions are always better

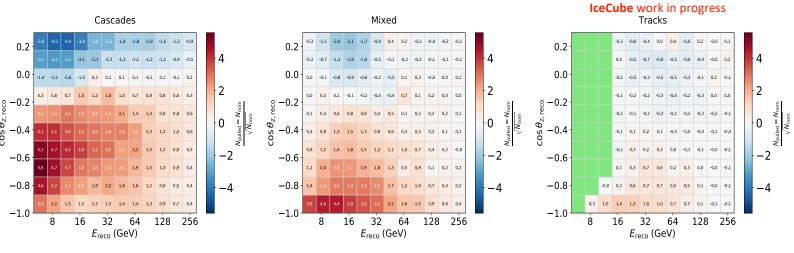




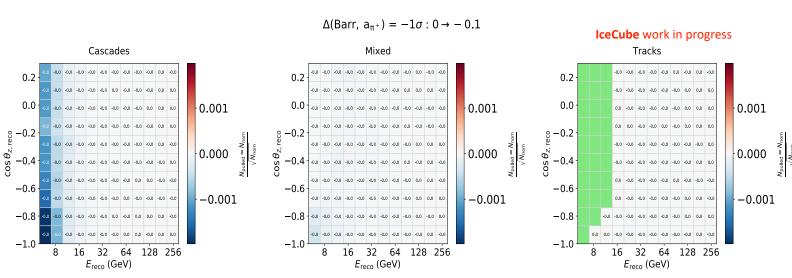
Systematics



- 30+ systematic uncertainties evaluated for upcoming analysis
 - Many have a negligible impact on oscillation measurements and are dropped
 - The non-negligible systematics can be <u>very</u> nonnegligible
 - e.g. DOM efficiency, hole ice photon scattering, spectral index
 - Subset of systematics checked for ν_{τ} appearance are in backup



 $\Delta(\text{Holeice}, p_0) = +0.5 : 0.102 \rightarrow 0.602$





Flux Uncertainties



*arXiv:astro-nh/0611266

- Using 'Barr blocks'* for hadroproduction fluxrelated uncertainties
 - ~18 total
- Matrix Cascade Equation (MCEq)** for flux
 - Fast and amenable to different cosmic ray spectra & hadronic interaction models
 - Correlated treatment of evolving flux particle interactions & decays

							ai Aiv.asti 0-	011/0011200
E _i (GeV)		Pic	Pions			Kaons		
<8	10%		30%				40%	
8-15	309	% 10%	3	0%		40%		
15-30	30	10 5%	1	0%		30 20	10%	
30-500	30	15%				40	30%	
>500	30	15%+Energy dep.				40	30%+Energy	dep.
0		0	.5	X _{LAB}	1 (0	0.5	\mathbf{X}_{LAB}

$$\frac{d\Phi}{dX} = [\hat{\Lambda}_{int} + \hat{\Lambda}_{dec}\rho^{-1}]\Phi$$

**arXiv:1503.00544

cross section uncertainty



- Relevant cross section uncertainties are included in analyses (DIS, RES, QE). Do we need an additional uncertainty for ν_{τ} specifically?
 - Independent of kinematic considerations (x, Q², W) or parton distribution functions, is a relative—or absolute—systematic uncertainty necessary for CC-DIS ν_{τ} , e.g.

$$r_{\tau/\mu} \equiv \sigma^{\text{CC}}(\nu_{\tau}N)/\sigma^{\text{CC}}(\nu_{\mu}N)$$
?

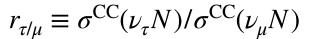
DIS cross section

$$\frac{\mathrm{d}^2 \sigma^{\nu/\bar{\nu}}}{\mathrm{d}x \mathrm{d}y} = \frac{G_F^2 M_N E_{\nu}}{\pi (1 + Q^2 / M_W^2)^2} \left\{ (y^2 x + \frac{m_l^2 y}{2E_{\nu} M_N}) F_1(x, Q^2) + \left[(1 - \frac{m_l^2}{4E_{\nu}^2}) - (1 + \frac{M_N x}{2E_{\nu}}) y \right] F_2(x, Q^2) \right\}$$

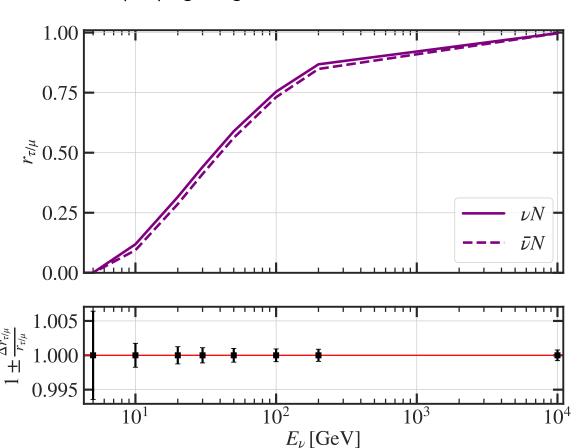
$$\pm \left[xy (1 - \frac{y}{2}) - \frac{m_l^2 y}{4E_{\nu} M_N} \right] F_3(x, Q^2) + \frac{m_l^2 (m_l^2 + Q^2)}{4E_{\nu}^2 M_N^2 x} F_4(x, Q^2) - \frac{m_l^2}{E_{\nu} M_N} F_5(x, Q^2) \right\}$$

The cross section ratio + uncertainties



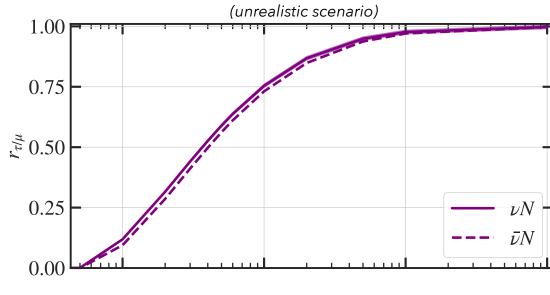


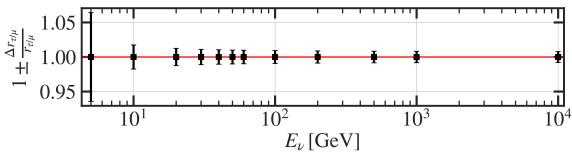




~0.5% relative uncertainty at 5 GeV → *small!*

propagating $10 \times \text{the } F_i \text{ uncertainties}$



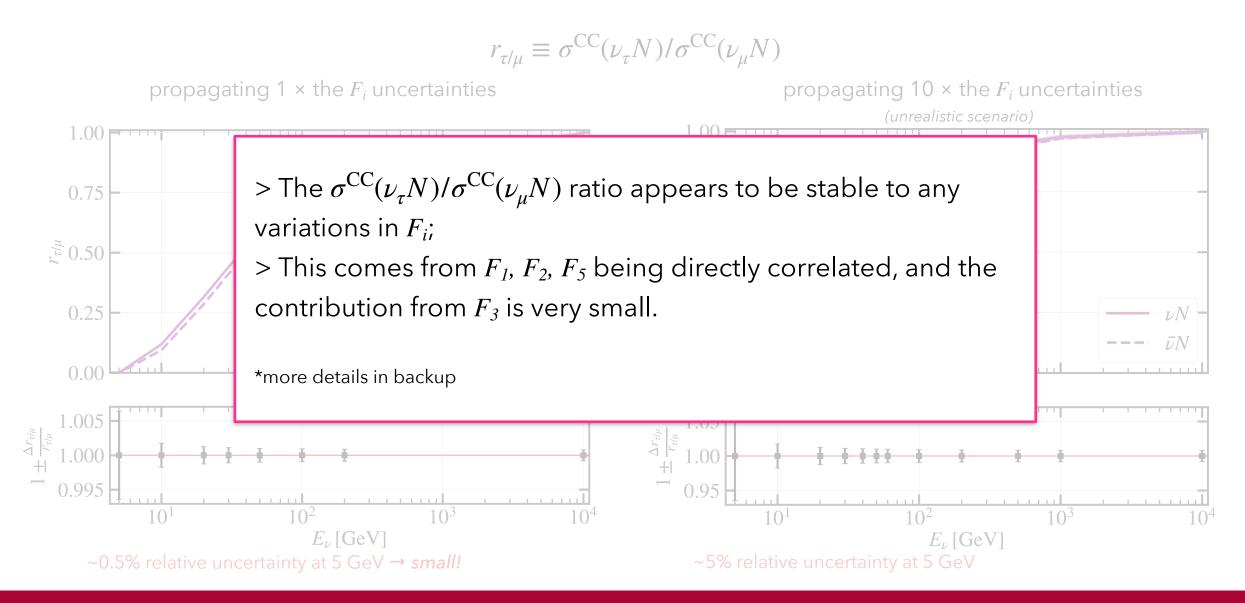


~5% relative uncertainty at 5 GeV

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The cross section ratio + uncertainties









THE FUTURE

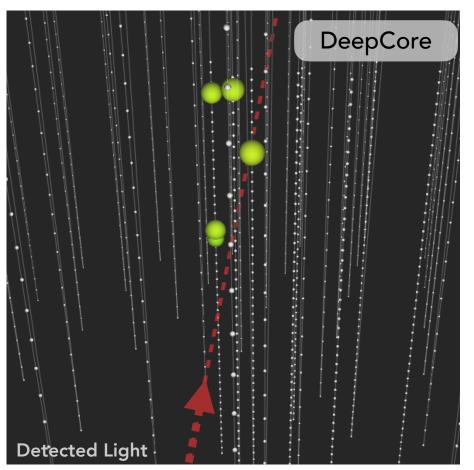


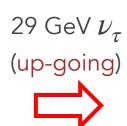
IceCube Upgrade

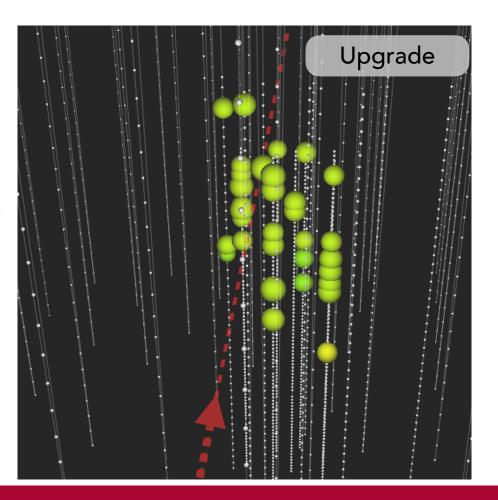


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Dense instrumentation within inner core



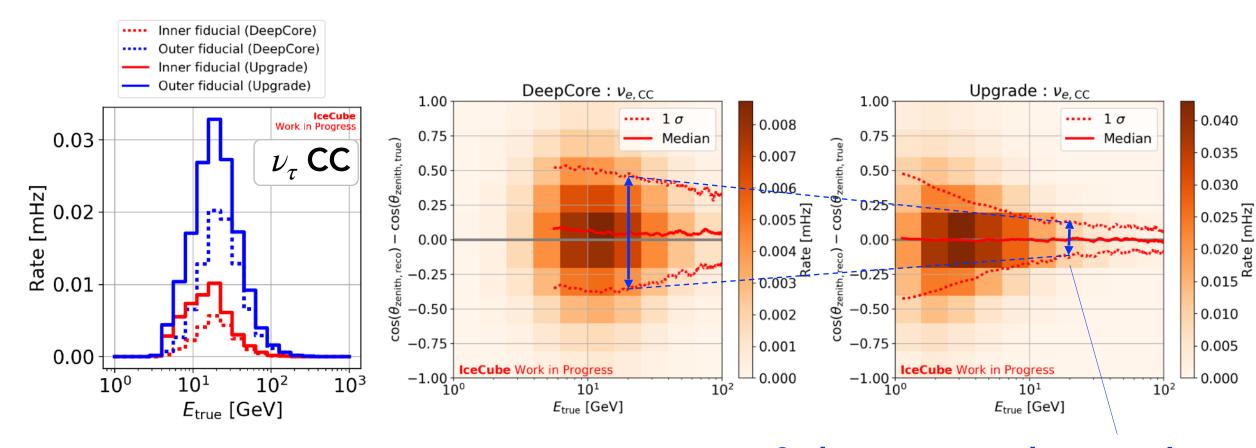






IceCube Upgrade





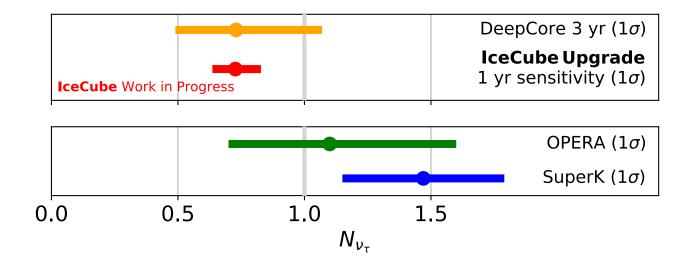
 $^*\nu_e$ CC and ν_τ CC both appear as cascades. ν_e are an easier proxy for cascade reconstruction development.

3x improvement in cascade resolution @ ν_{τ} appearance energies

IceCube Upgrade



- Conservative experimental choices still illustrate potential of IceCube Upgrade for physics
 - ~10% $N_{\nu_{\tau}}$ resolution with 1-year of data
 - Excludes improvements from new reconstructions, better detector systematics, better flux treatment, and no combination of 10+ years of DeepCore data



Conclusion



- Leading ν_{τ} appearance measurement(s) & constraints for (non)unitarity through $N_{\nu_{\tau}}$

• IceCube Upgrade provides incredible future for oscillation results, notably ν_{τ}





BACKUP

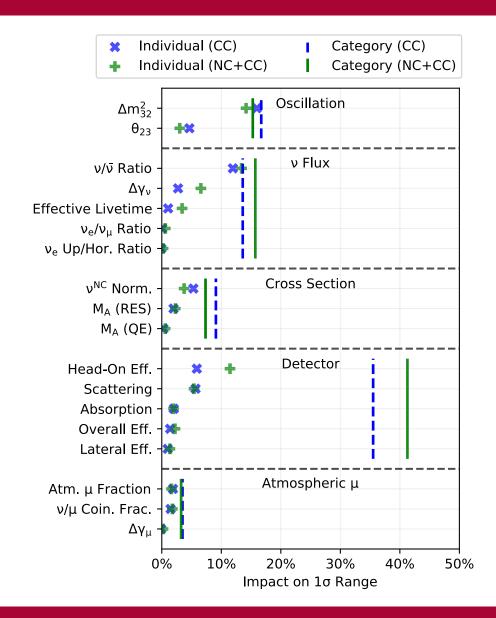


Systematic Uncertainties



• Figure is relative impact from 3-year ν_{τ} appearance analysis for the nonnegligible uncertainties

 Updated list for upcoming 8year analysis nearing finalization



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Outline of the ν_{τ} xsec relative uncertainty study

1. Take the good old Reno & Kretzer (2002) expression for the DIS cross section:

$$\frac{\mathrm{d}^2 \sigma^{\nu/\bar{\nu}}}{\mathrm{d}x \mathrm{d}y} = \frac{G_F^2 M_N E_{\nu}}{\pi (1 + Q^2 / M_W^2)^2} \left\{ (y^2 x + \frac{m_l^2 y}{2E_{\nu} M_N}) F_1(x, Q^2) + \left[(1 - \frac{m_l^2}{4E_{\nu}^2}) - (1 + \frac{M_N x}{2E_{\nu}}) y \right] F_2(x, Q^2) \right.$$

$$\pm \left[xy (1 - \frac{y}{2}) - \frac{m_l^2 y}{4E_{\nu} M_N} \right] F_3(x, Q^2) + \frac{m_l^2 (m_l^2 + Q^2)}{4E_{\nu}^2 M_N^2 x} F_4(x, Q^2) - \frac{m_l^2}{E_{\nu} M_N} F_5(x, Q^2) \right\}$$

- **2.** Compute the structure functions $F_i(x,Q^2)$:
 - 2.1. Express $F_i(x,Q^2)$ in terms of the parton distribution functions (PDFs) for individual quarks, $q(x,Q^2)$;
 - 2.2. Choose a PDF set and extract the respective $q(x,Q^2)$ ± uncertainty.

3.
$$q(x,Q^2) \pm \Delta q(x,Q^2) \longrightarrow F_i(x,Q^2) \pm \Delta F_i(x,Q^2) \longrightarrow \sigma_l(E) \pm \Delta \sigma_l(E) \longrightarrow \frac{\sigma_{\nu_\tau}}{\sigma_{\nu_\mu}}(E) \pm \Delta \left(\frac{\sigma_{\nu_\tau}}{\sigma_{\nu_\mu}}(E)\right)$$

*T. Kozynets, NBI

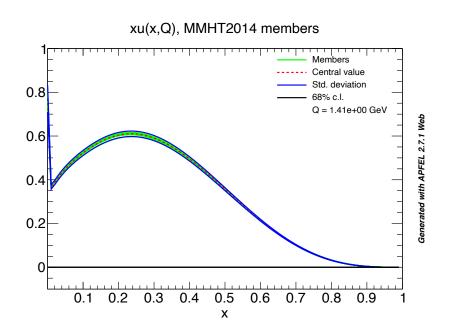
2.1. Computing the structure functions

		u(x), c(x) etc. are PDF-specific
proton target	$F_{2p}^{\nu}(x) = 2x[d(x) + s(x) + \bar{u}(x) + \bar{c}(x)]$ $xF_{3p}^{\nu}(x) = 2x[d(x) + s(x) - \bar{u}(x) - \bar{c}(x)]$	$F_{2p}^{\bar{\nu}}(x) = 2x[u(x) + c(x) + \bar{d}(x) + \bar{s}(x)]$ $xF_{3p}^{\bar{\nu}}(x) = 2x[u(x) + c(x) - \bar{d}(x) - \bar{s}(x)]$
neutron target	$F_{2n}^{\nu}(x) = 2x[u(x) + s(x) + \bar{d}(x) + \bar{c}(x)]$ $xF_{3n}^{\nu}(x) = 2x[u(x) + s(x) - \bar{d}(x) - \bar{c}(x)]$	$F_{2n}^{\bar{\nu}}(x) = 2x[d(x) + c(x) + \bar{u}(x) + \bar{s}(x)]$ $xF_{3n}^{\bar{\nu}}(x) = 2x[d(x) + c(x) - \bar{u}(x) - \bar{s}(x)]$
isoscalar target	$F_{iN} = \frac{F_{ip}}{}$	$+F_{in} \over 2$ \leftarrow assumed in this work
leading order relations Albright-Jarlskog;	$F_1(x) =$	$F_5(x) = \frac{F_2(x)}{2x}; F_4(x) = 0.$

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2.2. Accessing the individual quark PDFs

> The <u>LHAPDF</u> database provides the PDF sets with measurements & fits from different groups, for example:



LHAPDF 6.3.0



Official LHAPDF 6.3 PDF sets: currently 1045 available, of which 1044 are validated.

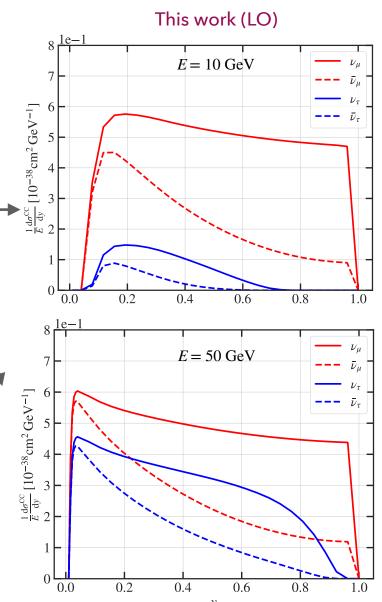
LHAPDF ID	Set name and links		Number of set members	Latest data version	Notes
251	GRVPI0 (tarball)	(info file)	1	1	alpha_s was broken in LHAPDF5. This version uses approximate 1st order running from reported Lambda4,5 values.
252	GRVPI1 (tarball)	(info file)	1	1	
270	xFitterPI_NLO_EIG (tarball)	(info file)	8	1	
280	xFitterPI_NLO_VAR (tarball)	(info file)	6	1	
10000	cteq6 (tarball)	(info file)	41	-1	Corresponds to LHAPDF5's cteq6 or cteq6mE; central member equivalent to cteq6m.
10042	cteq6l1 (tarball)	(info file)	1	4	
10150	cteq61 (tarball)	(info file)	41	1	
10550	cteq66 (tarball)	(info file)	45	1	

- > Several Python libraries (such as parton or pdfflow) can process these files directly and extract the PDFs for a needed flavor;
- > The central value is normally "member 0";
- > The confidence limits (68 or 90%, stated in the info file) are given in the other "members" of the PDF set.

3. Building up the cross sections: $d\sigma/dy$

> Plugging the F_i into the DIS cross section formula, we get the following differential cross sections:

> The shapes and the relative scaling between the flavours are ~in agreement with those from Reno (right panel).



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